Cycle 25 Update and Propagation

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Agenda

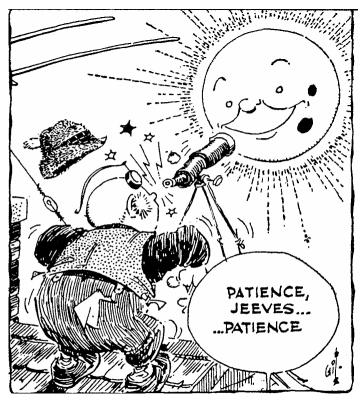
- Cycle 25
 - Historical data
 - Predictions
 - Latest data
- MF and HF propagation
 - General summary
 - Short-term opportunities
 - Some history
- Sporadic E propagation
 - Focus on 6m
 - Short-term probabilities
 - Long-term probabilities

EAA AirVenture 2022 – July 25-31

- My wife Vicky AE9YL and I attended AirVenture on Wednesday and Thursday
- Theme was the 75th anniversary of the USAF
 - Lots of fast and loud aircraft
- We assisted Bob Inderbitzen NQ1R (from HQ) at the ARRL booth
- Vicky visited the KidVenture area
 - EAA donated 500 of the ARRL receivers for kids to build
 - 40-150 MHz, \$14.95
- Stations on the air
 - W9ZL at KidVenture
 - W9W at ICOM
 - W1HQ at the ARRL booth (remoted to HQ)



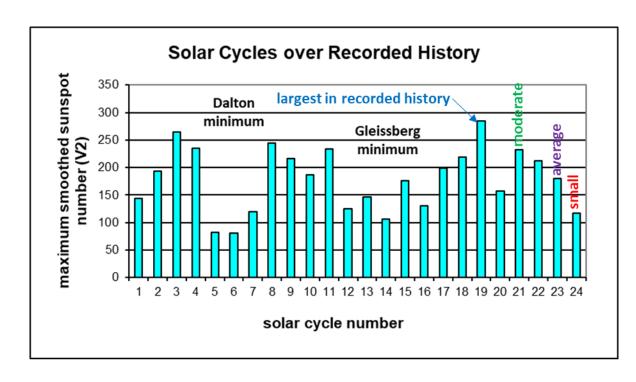
Cycle 25



from arrl.org

Historical Look at All 24 Cycles

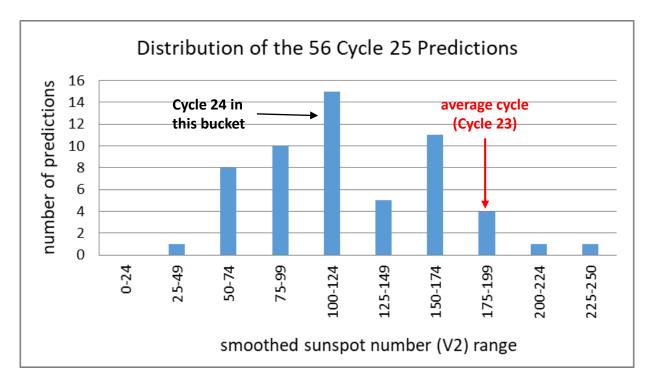
- Cycle 1 began in 1755
 - Maunder Minimum occurred from 1645-1715 with few sunspots
- We've gone through three periods of 'big' solar cycles
 - Cycles 1-4, 8-11, 17-23
- We've gone through two periods of 'small' solar cycles
 - Cycles 5-7, 12-16
- With Cycle 24, we appear to be in a third period of small solar cycles



Will Cycle 25 get us out of this third period of small solar cycles?

Let's look at predictions for Cycle 25

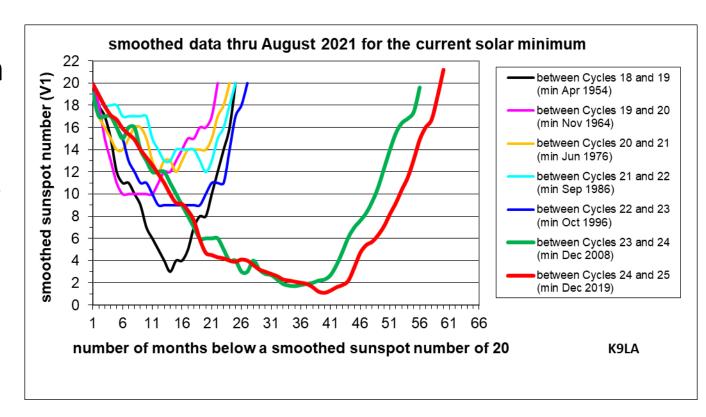
There Is a Consensus (but it doesn't mean it's correct!)



- 50 of the 56 predictions (89%) are for a below average cycle
- 4 are for an average cycle
- 2 are for a larger-than-average cycle

Solar Minimums in Our Lifetimes

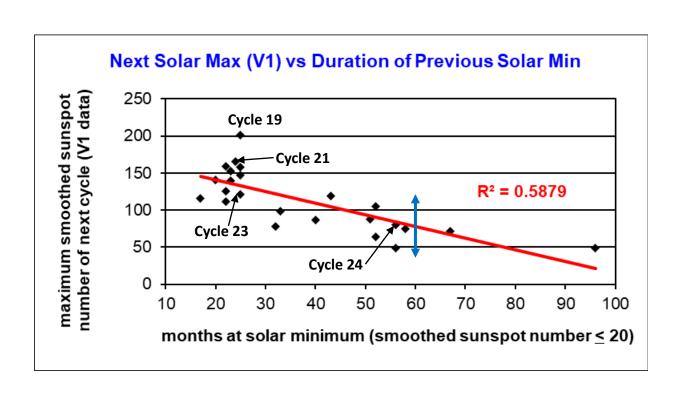
- We became used to short solar minimum periods – about 2 years duration
- But then along came the solar minimum between Cycles 23 & 24 – about 5 years duration
- Solar minimum between Cycles 24 & 25 was long, too



So what does this tell us?

Solar Minimum vs Next Solar Maximum

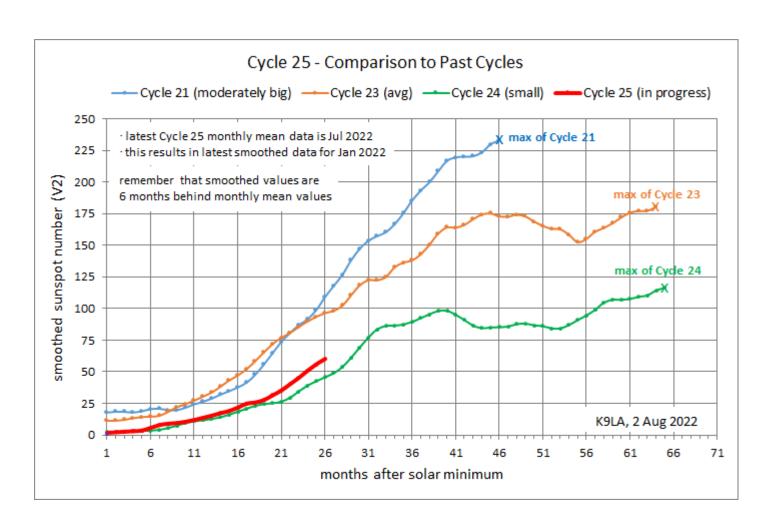
- The longer the solar minimum duration, the smaller the next cycle
- But it's not a perfect correlation
- Cycle 19 biggest cycle in recorded history
- Cycle 21 moderately big cycle
- Cycle 23 an average cycle
- Cycle 24 smallest cycle in our lifetimes – 4th smallest cycle in recorded history



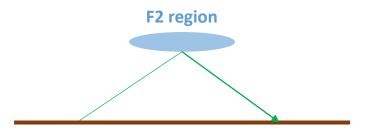
All we can do is wait to see what happens

How Is Cycle 25 Doing?

- All the curves start when the smoothed sunspot number numerically minimized
 - December 2019 for the start of Cycle 25
- We currently have 25 months of smoothed sunspot number data
- Where will Cycle 25 end up?
- Need more sunspots for worldwide 15m, 12m and 10m propagation via the F2 region on a daily basis



MF and HF Propagation

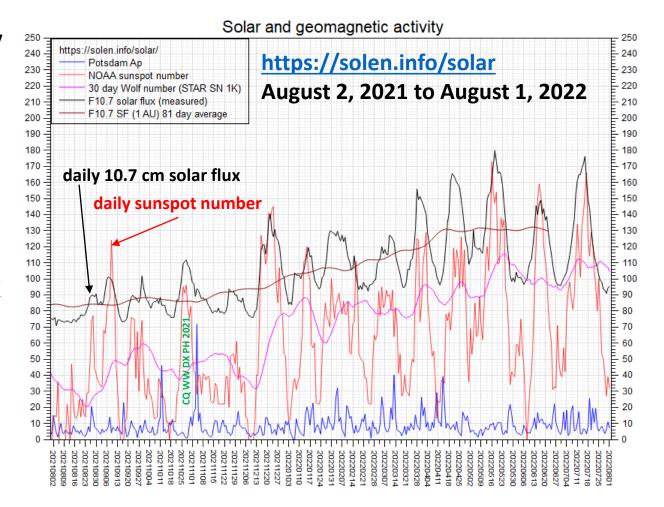


MF and HF Propagation

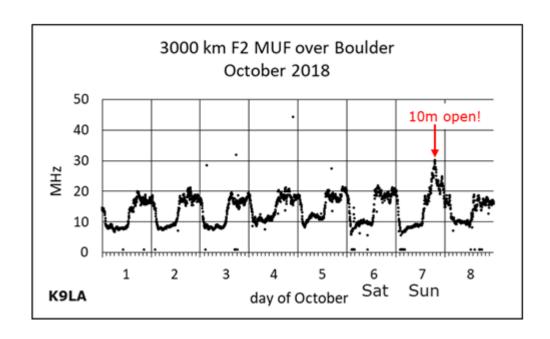
- 160m, 80m and 60m
 - Should be good at night unless there are thunderstorms
 - Might help to have a 'receive' antenna to minimize noise small loop
- 40m and 30m
 - Should be good during the day for shorter distance QSOs
 - Should be good at night for longer distance QSOs
- 20m and 17m
 - Should be good during the day
 - Enough sunspots yet for nighttime?
- 15m, 12m and 10m
 - Mostly north-south paths in the daytime
 - Should see good improvement this coming fall/winter
 - More east-west paths likely
 - Watch for sporadic E on 10m and 6m right now and in December

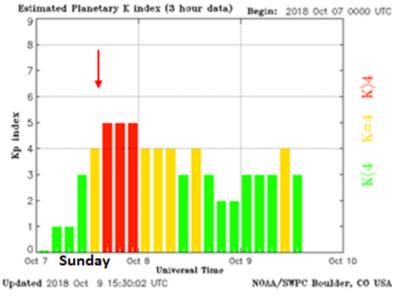
Short-Term Propagation Opportunities

- To reiterate, we still have a way to go before 15m, 12m and 10m will be open daily on a worldwide basis via the F2 region
- In the meantime, keep an eye on the daily sunspot number and the daily 10.7 cm solar flux
 - When they spike up, can give us short-term openings on these higher HF bands
 - Note CQ WW DX PH 2021 contest



An Example of a Spike in the K Index





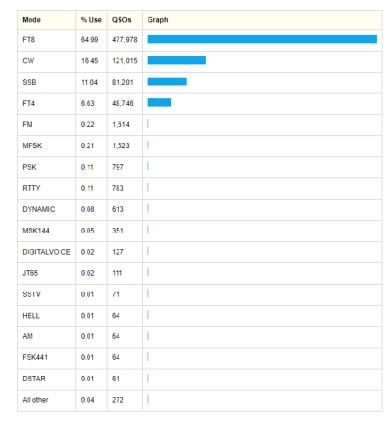
- 2018 California QSO Party
- No W6 stations heard on 10m in the Midwest on Saturday
- K index spiked up on Sunday lots of W6 stations to work

The Digital Modes

Weak-Signal S/N Limits

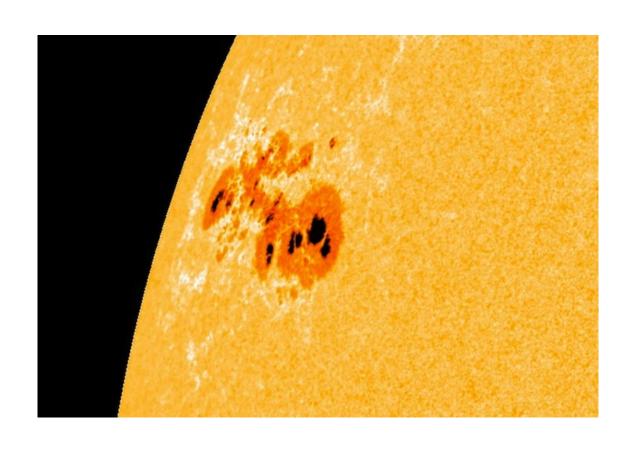
fastest data transfer at the top, slowest at the bottom Mode (B = 2500 Hz) SSB ~+10 dB MSK144 - 8 CW, "ear-and-brain" -15 -21 FT8 -23 JT4 JT65 -25 JT9 -27 QRA64 -27 WSPR -31

July 26 to Aug 2, 2022 (from Club Log data)



- Take advantage of the digital modes
- They made our last solar minimum the most active in history on the higher HF bands – and on 6m

Why Are Sunspots Important?



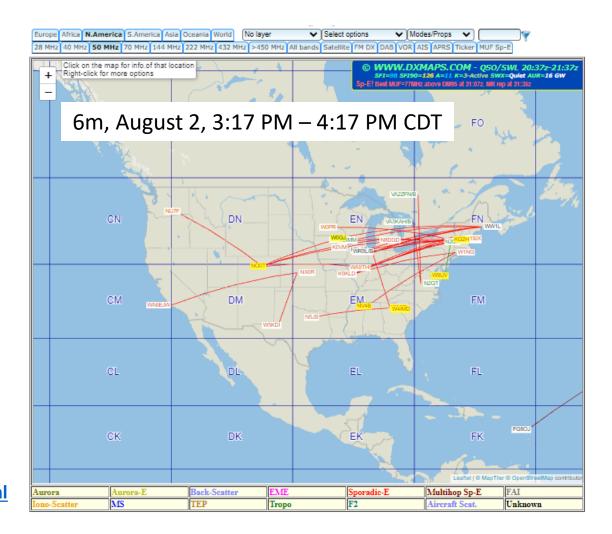
- The white area around a sunspot is called a plage (French for 'beach')
- These areas emit EUV radiation
- EUV radiation ionizes the atmosphere at F2 region altitudes
 - Sunspots themselves do not ionize anything (nor does 10.7 cm solar flux)
- F2 region is responsible for most of our long distance QSOs
 - And most QSOs at night

A Brief History of Solar Discoveries

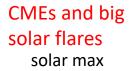
- Over 2000 years ago Chinese observed sunspots
- 1610 Galileo invented the telescope
- 1843 Schwabe credited with the discovery of the ~11 year cycle
- 1849 Wolf developed a method to count sunspots
- 1902 Kennelly (US) and Heaviside (U.K.) suggested independently that the Earth's upper atmosphere consisted of an electrically conductive region
- 1914 Hale credited with the discovery of the ~22 year cycle
- 1924 Appleton found conclusive evidence of an electrically conductive region by measuring the angle of arrival of radio waves from a nearby transmitter
- 1925 Breit and Tuve demonstrated the existence of the ionosphere with the first ionosonde
- 1930 Petit found a relationship between sunspots and EUV
- 1957-1958 International Geophysical Year (IGY) made worldwide measurements of the ionosphere
- 1970s first model of the ionosphere developed
- 1978 first propagation prediction program for Amateur Radio operators (MINI-MUF)

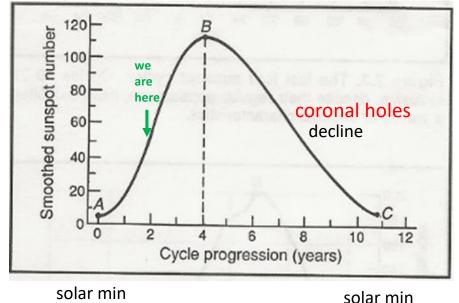
What Are the Bands Doing Right Now?

- Go to <u>dxmaps.com</u>
- Select a view
 - World, NA, Europe, . . .
- Select a band
 - 2200m to above 432 MHz
- Other methods
 - KC2G MUF map
 - https://prop.kc2g.com/
 - PSKReporter
 - https://pskreporter.info/pskmap.html
 - WSPRNet
 - https://www.wsprnet.org/drupal/
 - Reverse Beacon Network
 - http://www.reversebeacon.net/
 - IARU/NCDXF beacons
 - https://www.ncdxf.org/pages/beacons.html



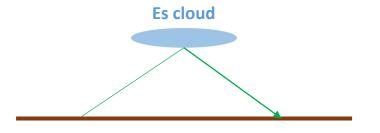
When Do Disturbances Occur?





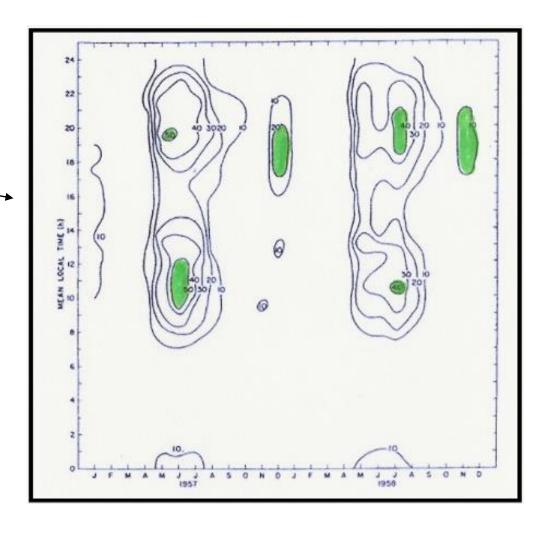
- We have to take the bad (CMEs and big solar flares) with the good (great propagation on the higher HF bands)
- CMEs most prevalent around solar max
 - Geomagnetic storms
- Big solar flares most prevalent around solar max
 - Solar radiation storms polar cap
 - Radio blackouts daylight side of Earth
- Coronal holes most prevalent during the decline of a solar cycle
- Quietest time is during the ascent of a solar cycle
 - Where we are right now

Sporadic E Propagation

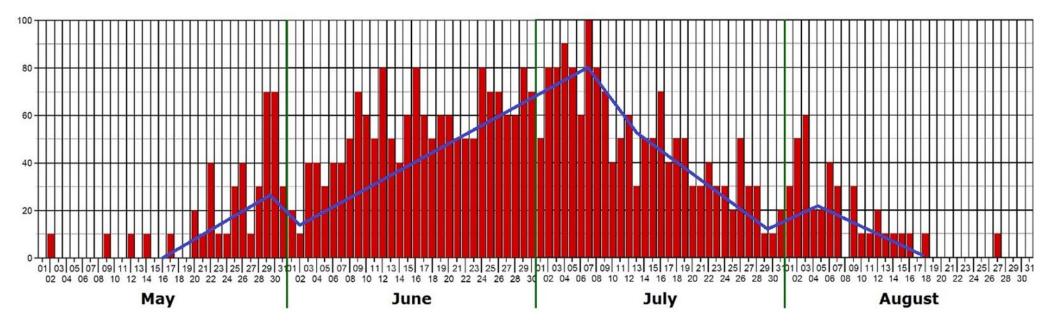


6-Meter Propagation (Short-Term)

- Sporadic-E (Es)
- Probability vs month and local time
 - Green areas are highest probability
- Best <u>summer</u> local times are
 - Late morning
 - Early evening
 - Always monitor in-between, too



VE3EN 6-Meter Es Data (Long-Term)



- 2004-2014 (11 years of data)
- Eastern-Western Europe to Eastern Half North America
 - USA 1/2/3/4/8/9 call areas + Canada VO1,VE1,VE2,VE3,VE9,VY2
- 5 or more contacts between multiple stations
- Y-axis is the % of the 11 years that QSOs were made
 - Early August has had 6m Es openings on about 3 of those 11 years

Summary

- Cycle 25 is in its ascent solar maximum around 2025
- So far Cycle 25 kind of looks like another small-ish cycle
 - The next 6-12 months should pin things down better
 - Hopefully it will get up to an 'average' cycle (like Cycle 23)
- Even if it is a small cycle, solar maximum should offer worldwide propagation on a daily basis with modest power and modest antennas on the higher HF bands and on 6-meters in the fall/winter months
- The digital modes offer an advantage over CW and SSB
 - Can decode a signal farther down in the noise
 - How much farther down depends on which digital mode
 - This is a significant advantage on 10m and 6m
- Take advantage of the summer Es season on 10m and 6m it's almost over
- There are tools on the internet to determine what the bands are doing right now – don't have to get deep into the SFI/SN/A, etc